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Small Fermi surfaces and strong correlation effects in Dirac materials with holography

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ABSTRACT: Recent discovery of transport anomaly in graphene demonstrated that a system known to be weakly interacting may become strongly correlated if system parameter (s) can be tuned such that fermi surface is sufficiently small. We study the strong correlation effects in the transport coefficients of Dirac materials doped with magnetic impurity under the magnetic field using holographic method. The experimental data of magneto-conductivity are well fit by our theory, however, not much data are available for other transports of Dirac material in such regime. Therefore, our results on heat transport, thermo-electric power and Nernst coefficients are left as predictions of holographic theory for generic Dirac materials in the vicinity of charge neutral point with possible surface gap. We give detailed look over each magneto-transport observable and 3Dplots to guide future experiments.

KEYWORDS: Holography and condensed matter physics (AdS/CMT), Black Holes in String Theory

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1 Introduction

Understanding strongly correlated electron systems has been a theoretical challenge for several decades [1]. Typically, excitations of such system lose particle nature, which invalidates Fermi-liquid theory, leaving physicists helpless. On the other hand, such systems also exhibit mysteriously rapid thermalization [2–5], which provides the hydrodynamic description [6, 7] near quantum critical point (QCP), where the system becomes universal: almost all details of the system are washed out. This is very analogous to the universality of a black hole in the sense that it also lose all the information of its mother star apart from the criticality index and a few conserved quantum numbers.

The gauge-gravity duality [8–10] provided a mathematically rigorous example and suggested a natural setting to put the analogy on more quantitative framework, which attracted much interest as a new paradigm for strongly interacting systems. More recently, large violation of Widermann-Frantz law was observed in graphene near charge neutral point, indicating that graphene is a strongly interacting system [11] in some windows of parameters. The gauge-gravity principle applied with two currents, exhibited remarkable agreement with the experimental data [12], improving the hydrodynamic analysis [7] of the same system. The fundamental reason for the appearance of the strong interaction in graphene is the *smallness of the fermi sea*: the effective coupling in a system with a Dirac cone is

$$\alpha_{\text{eff}} = \frac{e^2}{4\pi\epsilon\hbar c} \frac{c}{v_F} \sim 2.2/\epsilon_r,\tag{1.1}$$

with v_F fermi velocity. If the fermi surface passes near the Dirac point, the tip of the cone, electron hole pair creation from such small fermi surface is insufficient to screen the Coulomb interaction so that $\epsilon_r \sim 1$ and $\alpha_{\text{eff}} > 1$, making the system strongly interacting. This argument works even when a gap is open as far as the fermi surface can be tuned to be small. The eq. (1.1) also explains why electron-electron Coulomb interaction is small in usual metal where fermi surface is large. Since the above argument is so simple and generic, we expect that for any Dirac material, there should be a regime of parameters where electrons are strongly correlated. The presence of Dirac cone also provides reasoning why such system has a QCP with dynamical exponent z = 1 having Lorentz invariance.

The most well known Dirac material other than the graphene is the surface of a topological insulator (TI) [13, 14]. The latter has an unpaired Dirac cone and strong spin-orbit coupling, and as a consequence, it has a variety of interesting physics [15–17] including weak anti-localization (WAL) [18], quantized anomalous Hall effect [15], Majorana fermion [16] and topological magneto-electric effect [17]. Magnetic doping in TI can open a gap in the surface state by breaking the time reversal symmetry [19–21], and it is responsible for the transition from WAL to weak localization (WL). For extremely low doping, the sharp horn of the magneto-conductivity curve near zero magnetic field can be attributed to the particle nature of the basic excitations and indeed can be well described by Hikami-Larkin-Nagaoka (HLN) function [22]. However, for intermediate doping where the tendency of WAL and weak localization (WL) compete, a satisfactory theory is still wanted [19, 21, 23].¹

To understand why the transition regime is strongly interacting, look at the figure 1. We start with the case where fermi surface is large at zero doping. Increasing the surface gap pushes up the dispersion curve, which makes the fermi sea smaller. If gap is large enough but fermi surface is not small, the transition from WAL to WL happens in a manner involving sharp peak in MC curve, demonstrating the particle nature of the excitations. With more doping, the dispersion curve is pushed up more so that both large gap and small fermi surface are achieved and the system achieve the transition with strongly interacting nature. Therefore electron system near such transition region can be described by the holographic theory.

In the previous paper [27], we compared magneto-conductivity calculated by the holographic calculation and the data of Mn doped Bi_2Se_3 and Cr doped Bi_2Te_3 . We showed that the experimental data is well fit by the theoretical curves in the parameter island where fermi-surface is small. In this paper, we study all possible magneto-transport coefficients

¹For extremely thin film case, there is a phenomenological description. In refs. [19, 24], the authors assigned weights for two HLN functions of opposite sign by hand to fit the data. In graphene case, WAL-to-WL transition is better understood [25, 26] in terms of inter-valley scattering versus spin-orbit interaction. The parameter there corresponds to charge density which does not induces gap, while for TI case is the magnetic doping rate inducing surface gap. We believe that the physics involved is different.



Figure 1. Evolution of density of state. As we increase the doping and thereby the surface gap of the TI, the fermi surface gets smaller.

including thermal and thermo-electric transports of surface states of topological insulators in the regime of strong correlation. We will give 3D plots of each of them. Since not much data are available for heat transport or thermo-electric transports of Dirac material in such regime, our study can be regarded as predictions of holographic theory for generic Dirac materials in the vicinity of charge neutral point.²

2 Gravity dual of the surface of TI with magnetic doping

Although our target is general Dirac material not just for Topological Insulator (TI), we want to setup holographic formalism to describe the surface of it, which is one of the most well studied material with Dirac cone. Phenomenologically, we will be interested in magneto-transport of TI surface as a consequence of surface gap which is generated by the magnetic doping.

2.1 Holographic formulation of the surface state

We setup the holographic model by a sequence of reasonings.

- 1. The key feature of Topological bulk band is the presence of a surface *normalizable* zero-mode. It happens when the bulk band is inverted and one known mechanism for band inversion is large spin-orbit interaction. So considering boundary is crucial to discuss TI.
- 2. On the other hand, in Holographic theory, having both bulk and boundary of a physical system is very difficult, if not impossible, since the bulk of the physical system is already at the boundary of AdS space. In this situation, we have to 'carefully' delete either bulk or boundary for holographic description, depending on one's goal. Our main physical observable is low energy transport dynamics which happens only in the surface while bulk is a boring insulator as a gapped system.

 $^{^{2}}$ Our treatment can be applied for the case with surface gap as well as the case without gap as far as the system can be considered as a conductor.

Therefore, we delete the bulk part of TI together with the question 'how the system became topological insulator'. Our goal is to describe the surface phenomena knowing the system is already a TI. This means that we consider the case when the interaction do not destroy the inverted band structure of the bulk. In our case this is justified because we get strongly interacting system only at the surface by deforming the surface band only such that the fermi surface becomes small by use of the magnetic doping at the surface.

- 3. We focus on the consequence of the Dirac cone rather than the cause of the Dirac cone, the latter being the question of bulk. Once we confine our attention to the surface, we can characterize it by the presence of single Dirac cone. The surface physics due to the Dirac cone is not much different from that of graphene. Therefore, by treating the surface as a Dirac material, we already encoded the most important consequence of topological nature of the bulk band.
- 4. Our gravitational system is a deformation of charged AdS blackhole, which is widely used one. Our point is that one has to ask "for what material is such canonical gravity solution good?" Such local Lorentz invariant gravitation solutions are good only for Dirac materials, which, as a quantum critical system, can be characterized by the dynamical exponent z = 1.
- 5. There is one essential difference between a genuine 2 dimensional Dirac material and surface of TI. It is the relative position of the Fermi level and the surface band. For the former (like pure graphene), fermi level should pass the Dirac point at zero applied chemical potential. For the latter, it is not necessarily so. See figure attached. This point was also emphasized by Witten in his lectures on Topological material [28]. The position of fermi level is determined by the bulk physics, and there is no reason why it should pass the Dirac point of the surface band. This is the reason why the surface fermi energy at zero chemical potential is off the Dirac point of the surface Dirac cone. See the figure 1. This is what we mean 'delete carefully'.
- 6. The existence of the surface mode is the primary consequence of topological band structure, and whether surface gap is open or not is secondary question.
- 7. We want to discuss the transports of magnetically doped surface of TI. The question is what is the recipe to describe the system. We need the interaction between the magnetic impurity and the surface electron and this is the central part for the formulation for the practical purpose. Our interaction term should be the minimal one that describes the interaction with magnetic impurity with the charge current that breaks time reversal invariance.

2.2 Setup and background solutions

While it is clear that we need metric, gauge field and scalars to care the energy-momentum, the current and impurities respectively, for TI, special care is necessary to encode strong spin-orbit coupling (SOC). The latter induces the band inversion which in turn induces

massless fermions at the boundary and the topological nature of the system. To encode the effect of SOC in the presence of the magnetic impurities breaking the time reversal symmetry (TRS), we introduce a coupling between the impurity density and the instanton density. Such an interaction term was first introduced in [29] by us to discuss the SOC with TRS broken. It is the leading order term that can take care of gauge field coupling with impurity density in a TRS breaking manner.

With these preparations, our holographic model is defined by the Einstein-Maxwell-scalar action on an asymptotically AdS_4 manifold \mathcal{M} ,

$$2\kappa^{2}S = \int_{\mathcal{M}} d^{4}x \sqrt{-g} \left\{ R + \frac{6}{L^{2}} - \frac{1}{4}F^{2} - \sum_{I,a=1,2} \frac{1}{2} (\partial \chi_{I}^{(a)})^{2} \right\} - \frac{q_{\chi}}{16} \int_{\mathcal{M}} \sum_{I} (\partial \chi_{I}^{(2)})^{2} \epsilon^{\mu\nu\rho\sigma} F_{\mu\nu} F_{\rho\sigma} + S_{c}, \qquad (2.1)$$

$$S_c = -\int_{\partial \mathcal{M}} d^3x \sqrt{-\gamma} \left(2K + \frac{4}{L} + R[\gamma] - \sum_{I,a=1,2} \frac{L}{2} \nabla \chi_I^{(a)} \cdot \nabla \chi_I^{(a)} \right), \qquad (2.2)$$

where $\kappa^2 = 8\pi G$ and L is the AdS radius and we set $2\kappa^2 = L = 1$. S_c is the counter term for holographic renormalization. Here we introduce two scalar fields and only one scalar field contributes in the interaction term. The equations of motion are

$$\partial_{\mu} \left(\sqrt{-g} g^{\mu\nu} \sum_{a} \partial_{\nu} \chi_{I}^{(a)} \right) + \frac{q_{\chi}}{8} \partial_{\mu} (\epsilon^{\rho\sigma\lambda\gamma} F_{\rho\sigma} F_{\lambda\gamma} g^{\mu\nu} \partial_{\nu} \chi_{I}^{(2)}) = 0 \qquad (2.3)$$

$$\partial_{\mu} \left(\sqrt{-g} F^{\mu\nu} + \frac{q_{\chi}}{4} g^{\rho\sigma} \sum_{I} \partial_{\rho} \chi_{I}^{(2)} \partial_{\sigma} \chi_{I} \epsilon^{\alpha\beta\mu\nu} F_{\alpha\beta} \right) = 0 \qquad (2.4)$$

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu} \left[R + 6 - \frac{1}{4}F^2 - \sum_{I,a} \frac{1}{2}(\partial\chi_I^{(a)})^2 \right] - \frac{1}{2}F_{\mu}^{\rho}F_{\rho\nu} - \frac{1}{2}\sum_{I,a}(\partial_{\mu}\chi_I^{(a)})(\partial_{\nu}\chi_I^{(a)}) \\ - \frac{1}{\sqrt{-g}}\frac{q_{\chi}}{16}\sum_{I}(\partial_{\mu}\chi_I^{(2)})(\partial_{\nu}\chi_I^{(2)})\epsilon^{\rho\sigma\lambda\gamma}F_{\rho\sigma}F_{\lambda\gamma} = 0.$$
(2.5)

Assuming the solution of the equations of motion takes the form,

$$A = a(r)dt + \frac{1}{2}H(xdy - ydx),$$

$$\chi_{I}^{(1)} = \begin{pmatrix} \alpha x \\ \alpha y \end{pmatrix}, \qquad \chi_{I}^{(2)} = \begin{pmatrix} \lambda x \\ \lambda y \end{pmatrix},$$

$$ds^{2} = -\mathrm{U}(r)dt^{2} + \frac{dr^{2}}{\mathrm{U}(r)} + r^{2}(dx^{2} + dy^{2}),$$
(2.6)

we can find the exact solution as follows:

$$U(r) = r^2 - \frac{\alpha^2 + \lambda^2}{2} - \frac{m_0}{r} + \frac{q^2 + H^2}{4r^2} + \frac{\lambda^4 H^2 q_\chi^2}{20r^6} - \frac{\lambda^2 H q q_\chi}{6r^4},$$
(2.7)

$$a(r) = \mu - \frac{q}{r} + \frac{\lambda^2 H q_{\chi}}{3r^3},$$
(2.8)

where μ is the chemical potential, q and m_0 are determined by the conditions $A_t(r_0) = U(r_0) = 0$ at the black hole horizon $(r = r_0)$. q is the conserved U(1) charge and α and λ is relevant to momentum relaxation which will be discussed later:

$$q = \mu r_0 + \frac{1}{3}\theta H \qquad \text{with} \quad \theta = \frac{\lambda^2 q_{\chi}}{r_0^2}$$
(2.9)

$$m_0 = r_0^3 \left(1 + \frac{r_0^2 \mu^2 + H^2}{4r_0^4} - \frac{\alpha^2 + \lambda^2}{2r_0^2} \right) + \frac{\theta^2 H^2}{45r_0}.$$
 (2.10)

2.3 Thermodynamics and magnetisation

To obtain a thermodynamic potential for this black hole solution we compute the on-shell Euclidean action (S^E) by analytically continuing to Euclidean time (τ) of which period is the inverse temperature $t = -i\tau$, $S^E = -iS_{ren}$ where S^E is the Euclidean action. The temperature of the boundary system is identified by the Hawking temperature in the bulk,

$$4\pi T = U'(r_0) = 3r_0 - \frac{1}{4r_0^3} \left[H^2 + 2r_0^2(\alpha^2 + \lambda^2) + (q - H\theta)^2 \right], \qquad (2.11)$$

and the entropy density is given by the area of the horizon, $s = 4\pi r_0^2$. One can directly check that $dT/dr_0 > 0$ for positive r_0 , therefore, temperature is monotonically increasing function of r_0 . It implies that the entropy is monotonically increasing function of temperature.

From the Euclidean renormalized action, we can define the thermodynamic potential (Ω) and its density (W):

$$S^E \equiv \frac{\mathcal{V}_2}{T} \mathcal{W} \equiv \frac{\Omega}{T} \,, \tag{2.12}$$

where $\mathcal{V}_2 = \int dx dy$. Plugging the solution (2.7) into the Euclidean renormalised action, the potential density \mathcal{W} can be expresses as

$$\mathcal{W} = \frac{\Omega}{\mathcal{V}_2} = -m_0 + 2r_0^3 + \frac{(q^2 - H^2)}{2r_0} + \frac{q\theta H}{3r_0} - \frac{3\theta^2 H^2}{10r_0}.$$
 (2.13)

The boundary energy momentum tensor is given by

$$< T_{\mu\nu} > = \begin{pmatrix} 2m_0 & 0 & 0\\ 0 & m_0 & 0\\ 0 & 0 & m_0 \end{pmatrix},$$
 (2.14)

from which energy density $\varepsilon = 2m_0$. If we identify the boundary on-shell action to the negative pressure $-\mathcal{P}$ and combine background solutions, we get Smarr relation

$$\varepsilon + \mathcal{P} = s T + \mu q. \tag{2.15}$$

Variation of the potential density with respect to H, r_0 , μ and β gives

$$\delta \mathcal{W} = -\left(-\frac{H}{r_0} + \frac{q\,\theta}{3r_0} - \frac{\theta^2 H}{5r_0}\right)\delta H - s\delta T - q\delta\mu$$
$$-r_0\delta(\alpha^2) - \left(r_0 + \frac{q\,\theta H}{3r_0\lambda^2} - \frac{\theta^2 H^2}{5r_0\lambda^2}\right)\delta(\lambda^2),\tag{2.16}$$

where the variation of r_0 is replaced by that of temperature. If we define

$$\tilde{M} = -\frac{H}{r_0} + \frac{q\theta}{3r_0} - \frac{\theta^2 H}{5r_0}$$

$$\Theta_{\alpha} = r_0$$

$$\Theta_{\lambda} = r_0 + \frac{q\theta H}{3r_0\lambda^2} - \frac{\theta^2 H^2}{5r_0\lambda^2},$$
(2.17)

then (2.16) becomes

$$\delta \mathcal{W} = -\tilde{M}\delta H - s\delta T - q\delta\mu - \Theta_{\alpha}\delta(\alpha^2) - \Theta_{\lambda}\delta(\lambda^2).$$
(2.18)

Combining the variation of the second line in (2.13) and (2.16), we finally get the first law of thermodynamics;

$$\delta\varepsilon = -\tilde{M}\delta H + T\delta s + \mu\delta q - \Theta_{\alpha}\delta(\alpha^2) - \Theta_{\lambda}\delta(\lambda^2).$$
(2.19)

Now, we can define \tilde{M} as the magnetization of the 2+1 dimensional system from the theromorphic law. The magnetization \tilde{M} has finite value in the absence of the external magnetic filed.

$$\tilde{M}_0 \equiv \tilde{M}\Big|_{H=0} = \frac{q\theta}{3r_0} = \frac{q\lambda^2 q_{\chi}}{3r_0^3}.$$
(2.20)

We can interpret the boundary system as a ferro-magnetic material. The value of \tilde{M}_0 proportional to the charge density q and λ^2 . The scalar field χ_I plays role of the impurity density which cause momentum relaxation. From this analogy, we can identify λ^2 as magnetic impurity density and α^2 as non-magnetic one.

The thermodynamic law (2.19), contains variation of impurity denisty and the conjugate $\Theta_{\alpha,\lambda}$ can be interpreted as the energy dissipated per unit impurity.

The pressure also can be written as

$$\mathcal{P} = \langle T_{xx} \rangle + \tilde{M}H + \Theta_{\alpha}\alpha^{2} + \Theta_{\lambda}\lambda^{2}$$
$$= \frac{\varepsilon}{2} + \tilde{M}H + \Theta_{\alpha}\alpha^{2} + \Theta_{\lambda}\lambda^{2}.$$
(2.21)

The energy magnetization density can be defined as the linear response of the system under the metric fluctuation $\delta g_{tx}^{(0)} = B_E y$ [6, 30], that is, as the derivative of the on-shell action with respect to the B_E with temperature and chemical potential fixed.

$$\tilde{M}^{E} = \frac{1}{72r_{0}^{2}} \left[18q^{2}\theta + 3\theta H^{2}(2+\theta^{2}) - 4q H(9+4\theta^{2}) \right].$$
(2.22)

This result will be used when we calculate the DC transport coefficient in next section. It is straightforward to show that (2.22) is reduced to the energy magnetization of the dynomic black hole when we take $\theta \to 0$ limit.

We now discuss physical meaning of two impurities. As shown in (2.20), λ^2 plays role of the magnetic impurity density. On the other hand, in the absence of q_{χ} , impurity term in the thermodynamic first law (2.19) becomes $r_0(\alpha^2 + \lambda^2)$ which is same as the impurity density in non-magnetic theory [31]. Therefore, we can redefine the total impurity density (sum of magnetic and non-magnetic impurity density) and the ratio of the magnetic impurity density to the total impurity density as

$$\beta^2 = \alpha^2 + \lambda^2, \quad \gamma = \frac{\lambda^2}{\alpha^2 + \lambda^2}.$$
 (2.23)

From now on, we will use β^2 and γ instead of α and λ .

We finish this section with a comment on an issue on magnetization. In 2+1 dimensional system, there is a serious issue on the physical reality of magnetization, which is measured by looking the total magnetic induction B, which is the sum of external field and its effect induced inside the matter. However, the mass dimension of \tilde{M} and H are different and therefore we can not add these to form B. The presence of the problem is independent of using holography and we analyze the problem in appendix.

3 Magneto-transport coefficients

3.1 DC conductivities from horizon data

In this section, we calculate DC transports for the system with finite magnetization from black hole horizon data [32-34]. To do this, we turn on small fluctuations around background solution (2.7);

$$\delta G_{ti} = -t \mathbf{U}(r)\zeta_i + \delta g_{ti}(r)$$

$$\delta G_{ri} = r^2 \delta g_{ri}$$

$$\delta A_i = t(-E_i + \zeta_i a(r)) + \delta a_i(r),$$
(3.1)

where i = x, y and we also turn on the fluctuation of scalar fields $\delta \chi_i^{(a)}(r)$. With this ansatz equations of motion for fluctuation are time-independent. The linearized equations for the fluctuations are

$$\sum_{a} \delta \chi_{i}^{(a)''}(r) + \left(\frac{2}{r} + \frac{U'(r)}{U(r)}\right) \sum_{a} \delta \chi_{i}^{(a)}(r) - (\alpha + \lambda) \delta g'_{ri}(r) - (\alpha + \lambda) \left(\frac{2}{r} + \frac{U'(r)}{U(r)}\right) \delta g_{ri}(r) - \frac{(\alpha + \lambda)\zeta_{i}}{r^{2}U(r)} = 0,$$
(3.2)

$$\left[\mathbf{U}(r)\delta a_i' + a_t'(r)\delta g_{ti}(r) + \epsilon_{ij}H\mathbf{U}(r)\delta g_{ij}(r)\right]' + \epsilon_{ij}\left[\frac{q_{\chi}\lambda^2}{r^2}E_j + \left(\frac{H}{r^2} - \frac{2q_{\chi}\lambda^2a_t(r)}{r^3}\right)\zeta_j\right] = 0, \quad (3.3)$$

$$\begin{split} \delta g_{ti}''(r) &- \frac{1}{2r^4 \mathrm{U}(r)} \left[H^2 - 12r^4 + 2r^2 \beta^2 + 4r^2 \mathrm{U}(r) - r^4 a_t'(r)^2 + 4r^3 U'(r) + 2r^4 U''(r) \right] \delta g_{ti}(r) \\ &+ a_t'(r) \delta a_i(r) + \epsilon_{ij} H \left[a_t'(r) \delta g_{tj}(r) + \frac{a_t(r)}{r^2 \mathrm{U}(r)} \zeta_j - \frac{E_j}{r^2 \mathrm{U}(r)} \right] = 0 \\ \left[\frac{H^2}{4r^2} - 3r^2 + \frac{\beta^2}{2} - \frac{r^2 a_t'(r)^2}{4} + rU'(r) + \frac{r^2 U''(r)}{2} \right] \delta g_{ri}(x) - \frac{\lambda}{2} \left[1 - \frac{Hq_{\chi} a_t'(r)}{r^2} \right] \delta \chi_i^{(2)'}(r) \\ &- \frac{\alpha}{2} \delta \chi_i^{(1)} - \epsilon_{ij} \frac{H}{2r^2} \left[\delta a_j'(r) + \frac{a_t'(r) \delta g_{tj}(r)}{\mathrm{U}(r)} \right] + \frac{a_t'(r)}{2\mathrm{U}(r)} E_i - \left[\frac{1}{r} + \frac{a_t(r) a_t'(r)}{2\mathrm{U}(r)} - \frac{U'(r)}{2\mathrm{U}(r)} \right] \zeta_i = 0. \end{split}$$

$$(3.4)$$

(3.2) and (3.3) come from the equations from the scalars and the gauge field fluctuation and the last two equations (3.4) are the Einstein equations of the metric fluctuations.

To make the fluctuation equations to be regular at the horizon, we impose

$$\delta g_{ti} \sim \delta g_{ti}^{(0)} + \mathcal{O}(r - r_0)$$

$$\delta g_{ri} \sim \frac{\delta g_{ti}^{(0)}}{r^2 \mathcal{U}(r)} + \mathcal{O}(r - r0)$$

$$\delta \chi_i \sim \mathcal{O}(r - r_0)$$

$$\delta a_i \sim -\frac{E_i}{4\pi T} \log(r - r_0) + \mathcal{O}(r - r_0), \qquad (3.5)$$

where the last line implies in-falling condition of the gauge field fluctuation at the horizon. We can easily see this in the Eddington-Finkelstein coordinate.

With the regularity condition (3.5), the last equation can be written in terms of the black hole data and external sources as

$$\frac{\mathcal{F}}{r_0^2} \delta g_{tx}^{(0)} - \frac{H\mathcal{G}}{r_0^2} \delta g_{ty}^{(0)} + \mathcal{G}E_x + HE_y + 4\pi r_0^2 T\zeta_x = 0$$

$$\frac{\mathcal{F}}{r_0^2} \delta g_{ty}^{(0)} + \frac{H\mathcal{G}}{r_0^2} \delta g_{tx}^{(0)} + \mathcal{G}E_y - HE_x + 4\pi r_0^2 T\zeta_y = 0, \qquad (3.6)$$

where

$$\mathcal{F} = r_0^2 \beta^2 + (1+\theta^2) H^2 - q \,\theta H$$

$$\mathcal{G} = q - \theta H.$$
(3.7)

The solution of the algebraic equation (3.6) is

$$\delta g_{ti}^{(0)} = -\frac{r_0^2}{\mathcal{F}^2 + H^2 \mathcal{G}} \left[\mathcal{G}(\mathcal{F} - H^2) E_i + H(\mathcal{F} + \mathcal{G}^2) \epsilon_{ij} E_j + 4\pi r_0^2 T(\mathcal{F}\zeta_i + H\mathcal{G}\epsilon_{ij}\zeta_j) \right].$$
(3.8)

Now, let's consider current defined by

$$\begin{aligned} \mathcal{J}^{i} &= \sqrt{-g} F^{ir} \\ &= -\mathrm{U}(r) \delta a_{i}'(r) - a_{t}'(r) \delta g_{tx}(r) - \epsilon_{ij} \delta g_{rj}(r) \\ \mathcal{Q}^{i} &= \mathrm{U}(r)^{2} \partial_{r} \left(\frac{\delta g_{ti}(r)}{\mathrm{U}(r)} \right) - a_{t}(r) J^{i} \\ &= \mathrm{U}(r) \delta g_{ti}'(r) + (a_{t}(r) a_{t}'(r) - U'(r)) \delta g_{ti}(r) + a_{t}(r) \mathrm{U}(r) \delta a_{i}'(r) + \epsilon_{ij} H a_{t}(r) \mathrm{U}(r) \delta g_{rj}(r). \end{aligned}$$

$$(3.9)$$

One can easily show that these currents (3.9) become the electric current J^i and the heat current $Q^i = \langle T^{ti} \rangle -\mu J^i$ at the boundary $(r \to \infty)$. These quantities, however, can not be conserved along r direction in general. The way to resolve this problem is to take derivative of both side of (3.9) with respect to r. Substitute equations of motion (3.2), (3.3) and (3.4) to the right hand side and integrate it from black hole horizon to boundary. In summary;

$$J_{\text{total}}^{i} = \mathcal{J}(\infty) = \mathcal{J}(r_{0}) + \int_{r_{0}}^{\infty} dr \left(\partial_{r} \mathcal{J}(r) \Big|_{\text{e.o.m}} \right)$$
$$Q_{\text{total}}^{i} = \mathcal{Q}(\infty) = \mathcal{Q}(r_{0}) + \int_{r_{0}}^{\infty} dr \left(\partial_{r} \mathcal{Q}(r) \Big|_{\text{e.o.m}} \right), \qquad (3.10)$$

here, we put subscript 'total' to the current because it contains the magnetization current and the energy magnetization current contribution. By imposing the background solution and the horizon expansion (3.5), we get

$$J_{\text{total}}^{i} = E_{x} + \epsilon_{ij}\theta E_{j} - \frac{9}{r_{0}^{2}}\delta g_{ti}^{(0)} - \frac{H}{r_{0}}\epsilon_{ij}\delta g_{tj}^{(0)} + M\epsilon_{ij}\zeta_{j}$$
$$Q_{\text{total}}^{i} = -4\pi T\delta g_{ti}^{(0)} - M\epsilon_{ij}E_{j} - 2(M^{E} - \mu M)\epsilon_{ij}\zeta_{j}, \qquad (3.11)$$

where M and M^E are the magnetization and the energy mangetization defined in the previous section. The last term in the electric current and the last two terms in the heat current correspond to the mangetization current and the energy magnetization current, which should not be taken into account to the electric and heat current. Subtracting these terms and combining with the horizon data (3.8), we get the electric and heat current in terms of external sources;

$$J^{i} = \frac{(\mathcal{F} + \mathcal{G}^{2})(\mathcal{F} - H^{2})}{\mathcal{F}^{2} + H^{2}\mathcal{G}^{2}}E_{i} + \left[\theta + \frac{H\mathcal{G}(2\mathcal{F} + \mathcal{G}^{2} - H^{2})}{\mathcal{F}^{2} + H^{2}\mathcal{G}^{2}}\right]\epsilon_{ij}E_{j} + \frac{sT\mathcal{G}(\mathcal{F} - H^{2})}{\mathcal{F}^{2} + H^{2}\mathcal{G}^{2}}\zeta_{i} + \frac{sTH(\mathcal{F} + \mathcal{G}^{2})}{\mathcal{F}^{2} + H^{2}\mathcal{G}^{2}}\epsilon_{ij}\zeta_{j}$$
$$Q^{i} = \frac{sT\mathcal{G}(\mathcal{F} - H^{2})}{\mathcal{F}^{2} + H^{2}\mathcal{G}^{2}}E_{i} + \frac{sTH(\mathcal{F} + \mathcal{G}^{2})}{\mathcal{F}^{2} + H^{2}\mathcal{G}^{2}}\epsilon_{ij}E_{j} + \frac{s^{2}T^{2}\mathcal{F}}{\mathcal{F}^{2} + H^{2}\mathcal{G}^{2}}\zeta_{i} + \frac{s^{2}T^{2}H\mathcal{G}}{\mathcal{F}^{2} + H^{2}\mathcal{G}^{2}}\epsilon_{ij}\zeta_{j}. (3.12)$$

Now, the transport coefficients can be read off from

$$\begin{pmatrix} J^i \\ Q^i \end{pmatrix} = \begin{pmatrix} \sigma_{ij} & \alpha_{ij}T \\ \bar{\alpha}_{ij}T & \bar{\kappa}_{ij}T \end{pmatrix} \begin{pmatrix} E_j \\ -\frac{\nabla_j T}{T} \end{pmatrix},$$
(3.13)

where the temperature gradient $-(\nabla_i T)/T = \zeta_i$ in previous expression (3.12). The results are summarized as;

$$\sigma_{ii} = \frac{(\mathcal{F} + \mathcal{G}^2)(\mathcal{F} - H^2)}{\mathcal{F}^2 + H^2 \mathcal{G}^2}$$

$$\sigma_{ij} = \epsilon_{ij} \left[\theta + \frac{H\mathcal{G}(2\mathcal{F} + \mathcal{G}^2 - H^2)}{\mathcal{F}^2 + H^2 \mathcal{G}^2} \right]$$

$$\alpha_{ii} = \bar{\alpha}_{ii} = \frac{s\mathcal{G}(\mathcal{F} - H^2)}{\mathcal{F}^2 + H^2 \mathcal{G}^2}$$

$$\alpha_{ij} = \bar{\alpha}_{ij} = \epsilon_{ij} \frac{sH(\mathcal{F} + \mathcal{G}^2)}{\mathcal{F}^2 + H^2 \mathcal{G}^2}$$

$$\bar{\kappa}_{ii} = \frac{s^2 T \mathcal{F}}{\mathcal{F}^2 + H^2 \mathcal{G}^2}$$

$$\bar{\kappa}_{ij} = \epsilon_{ij} \frac{s^2 T H \mathcal{G}}{\mathcal{F}^2 + H^2 \mathcal{G}^2}.$$
(3.14)

The thermal conductivity when electric currents are set to be zero is

$$\kappa_{ij} = \bar{\kappa}_{ij} - T \,(\bar{\alpha} \cdot \sigma^{-1} \cdot \alpha)_{ij},\tag{3.15}$$

which can be calculated to be

$$\kappa_{ii} = \frac{s^2 T (\mathcal{F} + \mathcal{G}^2 - H^2 + 2H\mathcal{G}\theta + \mathcal{F}\theta^2)}{(\mathcal{F} + \mathcal{G}^2)^2 + H(H - 2\mathcal{G}\theta)(H^2 - 2\mathcal{F} - \mathcal{G}^2) + (\mathcal{F}^2 + H^2\mathcal{G}^2)\theta^2} \\ \kappa_{ij} = \epsilon_{ij} \frac{s^2 T (H - \mathcal{G}\theta)(\mathcal{G} + H\theta)}{(\mathcal{F} + \mathcal{G}^2)^2 + H(H - 2\mathcal{G}\theta)(H^2 - 2\mathcal{F} - \mathcal{G}^2) + (\mathcal{F}^2 + H^2\mathcal{G}^2)\theta^2}.$$
(3.16)

There are several interesting properties to the DC conductivities. The Onsaga's relation $\alpha_{ij} = \bar{\alpha}_{ij}$ can be easily checked. Also, notice that the off-diagonal components are anti-symmetric. These results are consistent with the DC transport coefficient of dyonic black hole [30, 35, 36] in $q_{\chi} \to 0$ or $\gamma \to 0$ limit, which gives some confidence for the validity of our results.

The other interesting observables are the Seebeck effect and the Nernst effect. Under the external magnetic field, electric field can be generated by the longitudinal thermal gradient. The generation of the longitudinal and the transverse electric field are called 'Seebeck' and 'Nernst' effect respectively. The Seebeck effect is used to check the Mott relation which is sensitive to the type of the interaction in material. Recently, people have found that the Mott relation near the charge neutrality point is broken in the clean graphene [37]. The Nernst effect is known as the phenomena of the vortex liquid and certain material shows large Nernst signal above critical temperature [38, 39]. The Seebeck coefficient (S) and the Nernst signal (N) are defined as

$$S = \left(\sigma^{-1} \cdot \alpha\right)_{xx}, \quad N = -\left(\sigma^{-1} \cdot \alpha\right)_{yx} \tag{3.17}$$

respectively and can be calculated from (3.14):

$$S = \frac{s(\mathcal{F} + \mathcal{G}^2)(\mathcal{G} + H\theta)}{(\mathcal{F} + \mathcal{G}^2)^2 + H(H - 2\mathcal{G}\theta)(H^2 - 2\mathcal{F} - \mathcal{G}^2) + (\mathcal{F}^2 + H^2\mathcal{G}^2)\theta^2}$$
$$N = \frac{s(\mathcal{F} - H^2)(H - \mathcal{G}\theta)}{(\mathcal{F} + \mathcal{G}^2)^2 + H(H - 2\mathcal{G}\theta)(H^2 - 2\mathcal{F} - \mathcal{G}^2) + (\mathcal{F}^2 + H^2\mathcal{G}^2)\theta^2}.$$
(3.18)

A technical remark is in order. Subtracting the contribution of the external magnetic field to the energy density amounts to shifting on-shell action to satisfy thermodynamic relations. It corresponds to adding the finite counter term $\frac{3H^2}{4r_0^2}$ to the action. We take the same procedure to calculate DC transports. The final electric and heat current (3.11) would be written in terms of \tilde{M} and \tilde{M}^E . After subtracting the magnetization current and the energy magnetization current contribution, we get same result of the DC transport coefficients (3.14).

3.2 Analysis of the transport coefficients

In this section, we analyze the external parameter dependence of magneto-transport coefficients. As shown in (3.14), DC transport coefficients are complicated function of the external parameters and hence we need full numerical calculation. In the absence of the external field, it gives anomalous transport which comes from non-zero magnetization of the system, which can be treated analytically. We will discuss several aspects of them.



Figure 2. Longitudinal conductivity σ_{xx} (a) as a function of Magnetic doping (γ) and magnetic field (B) at T = 2K, (b) as a function of Temperature and magnetic field at $q_{\chi}\gamma = 0.7$. In both case, we used $\beta^2 = \frac{2747}{(\mu m)^2}$ and $v_F = 7.5 \times 10^4 m/s$. We use the same number in all the figures appearing in this section.

3.2.1 Magnetotransport coefficients for non-ferromagnetic case ($\mu = 0$)

In the presence of the external magnetic field, DC transport coefficients (3.14) are complicated functions of other parameters. In particular we can not solve r_0 in terms of others but we can calculate transport coefficients numerically. In this section, we discuss transport coefficients in non-ferromagnetic case. As shown in (2.20), the magnetization at zero magnetic field proportional to the charge density or equivalently chemical potential. Therefore, we set $\mu = 0$ to discuss magneto-transports in non-ferromagnetic material.

Figure 2 shows the magnetic doping and temperature dependence of the longitudinal conductivity.

As discuss in [27], there is transition between weak-antilocalization and weak localization as we change magnetic doping or temperature. The longitudinal conductivity can be expressed at small magnetic field limit as

$$\sigma_{xx} \sim 1 + \frac{2}{r_0^2 \beta^2} (\frac{4}{9} \theta^2 - 1) H^2 + \mathcal{O}(H^4).$$
(3.19)

At high temprature, θ behaves $\sim 1/T^2$ and the longitudinal conductivity alway has negative curvature at H = 0. As temperature decreases, the sign of H^2 in (3.19) is flipped at $\theta = 3/2$. By combining with background solution, we get critical temperature given by the analytical expression

$$T_C = \frac{\sqrt{3/2}}{2\pi} \frac{(q_\chi \gamma - 1/4)}{\sqrt{q_\chi \gamma}} \beta.$$
(3.20)

If the value of $q_{\chi}\gamma$ smaller than 1/4, then T_C becomes negative and there is no transition to weak localization. The system has weak anti-localization in all temperature region. In the case of $q_{\chi}\gamma > 1/4$, the system shows weak localization for $T < T_C$ and transition to weak anti-localization appears at $T = T_C$. For high magnetic field, the longitudinal



Figure 3. Transverse conductivity σ_{xy} (a) as a function of Magnetic doping and magnetic field at T = 2K and (b) as a function of Temperature and magnetic field at $q_{\chi}\gamma = 0.7$.



Figure 4. longitudinal resistivity ρ_{xx} , (a) as a function of Magnetic doping and magnetic field at T = 2K and (b) as a function of Temperature and magnetic field at $\gamma = 0.5$.

conductivity becomes

$$\sigma_{xx}\Big|_{H\gg1} \sim \frac{\beta^2}{H} + \frac{2}{3} \left(\frac{q_{\chi}\beta^2\gamma}{H}\right)^2 + \mathcal{O}\left(\frac{1}{H^3}\right).$$
(3.21)

We see that total impurity term is dominant over magnetic impurity term.

Figure 3 shows transverse conductivity σ_{xy} as a function of magnetic doping, temperature and magnetic field. Here one can see that the σ_{xy} has maximum at H = 0 and the height is proportional to doping parameter and inverse temperature. We can understand this analytically since it can be expressed as

$$\sigma_{xy}\Big|_{H\ll 1} \sim \theta - \frac{4q_{\chi}\gamma}{3r_0^4}H^2 + \mathcal{O}(H^3)$$

$$\sigma_{xy}\Big|_{H\gg 1} \sim \frac{q_{\chi}\beta^2\gamma}{3H} + \mathcal{O}\left(\frac{1}{H^3}\right), \qquad (3.22)$$

here we use $r_0 \sim \sqrt{H}$ for $H \gg 1$.

The longitudinal resistivity is presented in figure 4. figure 4(a) shows magnetic doping dependence of magneto-resistance. One should notice that the signal of transition to WL



Figure 5. (a) Metalicity, the sign of $\partial \rho_{xx}/\partial T$, as a function of Magnetic doping and magnetic field. Negative (white) regions indicate the instability.

does not appear in the resistivity contrary to σ_{xx} shown in figure 2(a). One can understand it in the following way. The longitudinal resistivity comes from the inversion of the conductivity matrix as

$$\rho_{xx} = \frac{\sigma_{xx}}{\sigma_{xx}^2 + \sigma_{xy}^2},\tag{3.23}$$

where we use $\sigma_{xx} = \sigma_{yy}$ and $\sigma_{xy} = -\sigma_{yx}$. The relation $\rho_{xx} \sim 1/\sigma_{xx}$ holds only when σ_{xy} is sufficiently small. On the other hand, the weak localization in σ_{xx} appears in large doping region, where the transverse conductivity is large as one can see figure 3(a). Therefore, transverse component of the conductivity affects the longitudinal magneto-resistance and wash out the signal of weak localization. So it is better to define the weak localization by magneto-conductance instead of magneto-resistance at least for strongly interacting system. Figure 4(b) shows magneto-resistivity as function of (B,T) at a fixed doping rate. Near zero temperature, we can define the metallicity as the sign of $\partial \rho_{xx}/\partial T$, which is demonstrated at figure 5.

Due to the non vanishing θ , the transverse resistivity also has non-trivial behavior for other parameters. Near zero magnetic field and zero temperature, the transverse resistivity can be expressed as

$$\rho_{yx}\Big|_{T\ll 1, H\ll 1} \sim \frac{\tilde{\theta}\left(4\tilde{\theta}^4 + 13\tilde{\theta}^2 + 3\right)}{3\beta^4 \left(\tilde{\theta}^2 + 1\right)^2} H^2 + \mathcal{O}(H^4), \qquad (3.24)$$

where $\hat{\theta} = 6q_{\chi}\gamma$. On the other hand, the transverse resistivity linearly increases at large magnetic field limit as

$$\rho_{yx}\Big|_{H\gg 1} \sim \frac{8\tilde{\theta}}{45\beta^2} \cdot H + \cdots . \tag{3.25}$$

This behavior of the transverse resistivity is drawn in figure 6(a).



Figure 6. Transverse resistivity ρ_{yx} , (a) as a function of Magnetic doping and magnetic field at T = 0.01. (b) as a function of Temperature and magnetic field at $q_{\chi}\gamma = 0.7$.



Figure 7. Longitudinal thermal conductivity κ_{xx} (a) as a function of Magnetic doping and magnetic field at T = 2K, (b) as a function of Temperature and magnetic field at $q_{\chi}\gamma = 0.7$.

The impurity density dependence of the longitudinal thermal conductivity κ_{xx} also has non trivial behavior. For low temperature limit, κ_{xx} can be expanded as

$$\kappa_{xx}\Big|_{T\ll 1} \sim -\frac{8\pi^2 T (8\tilde{\theta}^4 + \tilde{\theta}^2 - 9)}{9\beta^4 (\tilde{\theta}^2 + 1)} H^2 + \mathcal{O}(H^4), \tag{3.26}$$

where $\tilde{\theta} = 6q_{\chi}\gamma$. Notice that the numerator of (3.26) changes sign as $\gamma(\text{or }\tilde{\theta})$ increases and it becomes zero at critical value $\tilde{\theta} = 1$. The dependences of κ_{xx} and κ_{yx} on impurity density γ , magnetic field *B* and temperature *T* are presented at figure 7 and figure 8 respectively. For given impurity density, κ_{xx} has power behavior in temperature as

$$\kappa_{xx}\Big|_{T\ll 1} \sim \frac{8\pi^2 T}{3} + \cdots$$

$$\kappa_{xx}\Big|_{T\gg 1} \sim \frac{64\pi^4 T^3}{9\beta^2} + \cdots$$
(3.27)

at zero magnetic field limit.



Figure 8. Transverse thermal conductivity κ_{xy} (a) as a function of Magnetic doping and magnetic field at T = 2K, (b) as a function of Temperature and magnetic field at q_{χ} .



Figure 9. Seebeck coefficient S (a) as a function of Magnetic doping and magnetic field at T = 2K, (b) as a function of Temperature and magnetic field at $q_{\chi}\gamma = 0.7$.

The Seebeck coefficient and the Nernst signal (3.18) is expanded near zero magnetic field as

$$S\Big|_{H\ll 1} \sim \frac{4\pi\theta}{3\beta^2(1+\theta^2)} \cdot H + \cdots$$

$$N\Big|_{H\ll 1} \sim \frac{4\pi(3+2\theta^2)}{3\beta^2(1+\theta^2)} \cdot H + \cdots$$
(3.28)

The exact calculations for Seebeck coefficient are drawn in figure 9 and figure 10.

3.2.2 Zero field limit and anomalous Hall transports

Taking the $H \to 0$ limit, the transport coefficients (3.14) become

$$\sigma = \begin{pmatrix} 1 + \frac{q^2}{r_0^2 \beta^2} & \theta \\ -\theta & 1 + \frac{q^2}{r_0^2 \beta^2} \end{pmatrix}, \quad \alpha = \begin{pmatrix} \frac{4\pi q}{\beta^2} & 0 \\ 0 & \frac{4\pi q}{\beta^2} \end{pmatrix}, \quad \bar{\kappa} = \begin{pmatrix} \frac{4\pi sT}{\beta^2} & 0 \\ 0 & \frac{4\pi sT}{\beta^2} \end{pmatrix}.$$
(3.29)

Notice that in the absence of the external field, only θ appears in σ_{xy} . The other transport coefficients are the same as those of the RN-AdS black hole with momentum relaxation.



Figure 10. Nernst signal N (a) as a function of Magnetic doping and magnetic field at T = 2K, (b) as a function of Temperature and magnetic field at $q_{\chi}\gamma = 0.7$.



Figure 11. The temperature and the charge density dependence of (a) the longitudinal conductivity σ_{xx} and (b) the transverse conductivity σ_{xy} .

Notice that there are no off-diagonal elements of α and $\bar{\kappa}$. In this zero magnetic field limit, the black hole horizon radius r_0 is independent of q_{χ} and it behaves as

$$r_0\Big|_{q\to\infty} \sim \frac{1}{3^{1/4}\sqrt{2}}\sqrt{q} + \mathcal{O}(1/\sqrt{q}), \qquad r_0\Big|_{T\to\infty} \sim \frac{4\pi}{3}T + \mathcal{O}(1/T).$$
 (3.30)

The longitudinal and the transverse conductivity are drawn in figure 11. The longitudinal conductivity σ_{xx} does not have q_{χ} dependence and hence is the same as one in RN-AdS with momentum relaxation. From (3.30), we can expand the longitudinal conductivity in small and large density region:

$$\sigma_{xx}\Big|_{q\ll 1} \sim 1 + \frac{36q^2}{(4\pi T + \sqrt{(4\pi T)^2 + 6\beta^2})^2 \beta^2} + \mathcal{O}(q^3)$$

$$\sigma_{xx}\Big|_{q\gg 1} \sim 2\sqrt{3}\frac{q}{\beta^2} - \frac{4\sqrt{2}T}{3^{1/4}\beta^2}\sqrt{q} + \mathcal{O}(1/\sqrt{q}).$$
(3.31)

In the limit of $\beta \gg T$, (3.31) becomes simpler form

$$\sigma_{xx}\Big|_{q\ll 1} \sim 1 + \frac{6q^2}{\beta^4}, \qquad \sigma_{xx}\Big|_{q\gg 1} \sim 2\sqrt{3}\frac{q}{\beta^2}.$$
(3.32)

In large carrier density limit, the longitudinal conductivity is linearly increasing in q, which is observed in the graphene. By solving Boltzman equation, it was shown to be [40-42]

$$\sigma_{xx} = \frac{e^2}{\hbar} \frac{(\hbar v_F)^2}{u_0^2} n,$$
(3.33)

where n is the charge carrier density which is the same as q in this paper and u_0^2 is proportional to the impurity density. If we identify $\frac{(\hbar v_F)^2}{u_0^2} \equiv \frac{2\sqrt{3}}{\beta^2}$, then the second line of (3.32) looks like the result of Fermi liquid theory. On the other hand, the longitudinal conductivity has quadratic behavior near charge neutrality point. If we introduce dimensionless quantity $\tilde{Q} \equiv 2\sqrt{3} q/\beta^2$, then

$$\sigma_{xx}\Big|_{q\ll 1} \sim 1 + \frac{1}{2}\tilde{Q}^2, \qquad \sigma_{xx}\Big|_{q\gg 1} \sim \tilde{Q}.$$
(3.34)

The transverse conductivity σ_{xy} is shown in figure 11(b). σ_{xy} has maximum at zero temperature and charge neutrality point and the height is $6q_{\chi}\gamma$. Large density and temperature behavior can be obtained from (3.30)

$$\sigma_{xy}\Big|_{q\gg 1} \sim \frac{q_{\chi}\beta^2\gamma}{q}, \qquad \sigma_{xy}\Big|_{T\gg 1} \sim \frac{q_{\chi}\beta^2\gamma}{T^2}.$$
(3.35)

The resistivity can be obtained by inverting the electric conductivity martix;

$$\rho_{xx} = \frac{\left(1 + \frac{q^2}{r_0^2 \beta^2}\right)}{\left(1 + \frac{q^2}{r_0^2 \beta^2}\right)^2 + \theta^2}, \qquad \rho_{xy} = -\frac{\theta}{\left(1 + \frac{q^2}{r_0^2 \beta^2}\right)^2 + \theta^2}.$$
(3.36)

Notice ρ_{xx} contains different information of σ_{xx} due to the presence of θ even in the absence of the external magnetic field. The transverse resistivity is proportional to θ which is related to the value of the magnetization. This phenomena is called anomalous Hall effect which comes from the intrinsic magnetic property of material. The scaling property of the anomalous Hall effect was discussed in [36]. In this paper, we focus on the charge and the temperature dependence of the resistivity. The effect of the magnetic impurity on the longitudinal resistivity is drawn in figure 12.

The figure 12(a) and (b) show the density and the temperature dependence of the longitudinal resistivity. In the absence of the magnetic impurity, the resistivity has maximum at the charge neutrality point in all temperature range. It is natural that due to the absence of charge carrier density at the charge neutrality point, the resistivity has maximum. On the other hand, in the presence of the magnetic impurity, the resistivity at the charge neutral point is suppressed at low temperature. As doping parameter γ increases, the suppressed region becomes larger. Figure 12(c) shows the temperature dependence of the longitudinal resistivity at the charge neutrality point for different value of γ . The density dependence of the longitudinal resistivity at low temperature is more interesting, see figure 12(d). In this figure, the maximum of the resistivity is not located at the charge neutrality point as γ increases. It can be understood as an effect of the magnetization which is proportional to θ . The denominator of the longitudinal resistivity (3.36) is maximized when θ has maximum



Figure 12. The charge and the temperature dependence of the longitudinal resistivity without (a) and with (b) magnetic impurity. (c) The temperature dependence of ρ_{xx} at the charge neutral point and (d) The charge density dependence of ρ_{xx} at low temperature (T = 2K) with different value of $q_{\chi}\gamma$.

value and it happens at the charge neutrality point as shown in figure 11(b). And the competition between the longitudinal and the transverse conductivity shift the maximum of the resistivity away from the charge neutrality point.

The effect of the magnetic impurity on the transverse resistivity is drawn in figure 13. As shown in (3.36), the transverse resistivity is proportional to θ and hence there is no transverse resistivity in the absence of the magnetic impurity. If we put magnetic impurity, there is maximum of the transverse resistivity at finite temperature and charge neutrality point, see figure 13(b). As we increase impurity density, the peak of transverse resistivity moves to high temperature region.

Non-zero value of σ_{xy} gives non-trivial θ dependence to the thermal conductivity. κ is defined by the thermal current without electric current as it was given by eq. (3.15). From (3.29), the longitudinal and the transverse thermal conductivity are

$$\kappa_{xx} = \frac{s^2 T}{r_0^2 \beta^2} \cdot \frac{\sigma_{xx} + \sigma_{xy}^2}{\sigma_{xx}^2 + \sigma_{xy}^2}$$

$$\kappa_{xy} = \frac{s^2 T q^2}{r_0^4 \beta^4} \cdot \frac{\sigma_{xy}}{\sigma_{xx}^2 + \sigma_{xy}^2}.$$
(3.37)

These results are shown in figure 14. Here, figure 14(a), (b) show κ_{xx} without and with magnetic impurity. In the absence of magnetic impurity, the longitudinal component of the thermal conductivity has maximum at charge neutrality point and grows as T at



Figure 13. The charge and the temperature dependence of the transverse resistivity without (a) and with (b) magnetic impurity. (c) The temperature dependence of ρ_{yx} at the charge neutrality point. (d) The charge density dependence of ρ_{yx} at low temperature (T = 2K) with different value of $q_{\chi}\gamma$.

low temperature and T^3 at high temperature. For finite magnetic impurity density, small dip appears at charge neutrality point. Figure 14(c) and (d) show the transverse thermal conductivity which shows q^2 behavior near charge neutrality point.

The Lorentz ratio is defined by the ratio between the longitudinal thermal conductivity and the longitudinal electric conductivity;

$$L \equiv \frac{\kappa_{xx}}{\sigma_{xx}T}.$$
(3.38)

In the large q limit, the transverse electric conductivity can be ignored and the Lorentz ratio is suppressed as 1/q

$$L\Big|_{q\gg1} \sim \frac{2\pi^2}{3\sqrt{3}} \cdot \frac{\beta^2}{q}.$$
(3.39)

On the other hand, at the charge neutrality point (q = 0), the longitudinal electric conductivity becomes 1 and the Lorentz ratio becomes

$$L\Big|_{q=0} \sim \left\{ \begin{array}{c} 8 \cdot \frac{\pi^2}{3} \ (T \ll 1) \\ \frac{(4\pi)^4}{9} \cdot \frac{T^2}{\beta^2} \ (T \gg 1) \end{array} \right\},\tag{3.40}$$

where we use

$$r_0 = \frac{2\pi}{3} \left(T + \sqrt{T^2 + 6\left(\frac{\beta}{4\pi}\right)^2} \right) \tag{3.41}$$



Figure 14. The longitudinal thermal conductivity κ_{xx} as a function of the charge and the temperature (a) without and (b) with magnetic impurity. (c) and (d) are Transverse thermal conductivity κ_{xy} without and with magnetic impurity respectively.



Figure 15. (a) Effect of interaction strength q_{χ} on Lorentz ratio near zero temperature. (b) Effect of impurity density on Lorentz ratio, $L^* = L \cdot \left(\frac{9}{(4\pi)^4 T^2}\right)$, at high temperature.

at the charge neutrality point. Figure 15 shows the charge dependence of Lorentz ration for various interaction strength q_{χ} and total impurity density β . In the figure, the violation of the Wiedemann-Frantz law is maximized at charge neutrality point and the violation region increases as interaction strength q_{χ} increases, figure 15(a) and the violation is suppressed as the total impurity density increases, figure 15(b). Notice that for large violation of the Wiedemann-Frantz law, one should use clean material ($\beta \ll 1$).



Figure 16. The temperature and the charge density dependence of Seebeck coefficient without (a) with (b) magnetic impurity.

Seebeck coefficient (S) and the Nernst signal (N) (3.17) can be written as

$$S = \frac{4\pi r_0^2 q (q^2 + r_0^2 \beta^2)}{(q^2 + r_0^2 \beta^2)^2 + (r_0^2 \beta^2)^2 \theta^2}$$
$$N = -\frac{4\pi r_0^4 q \beta^2 \theta}{(q^2 + r_0^2 \beta^2)^2 + (r_0^2 \beta^2)^2 \theta^2},$$
(3.42)

where we use (3.29). Notice that the Nernst signal is non zero in the absence of the external magnetic field because of the non zero component of the transverse electric conductivity. Both of S and N are odd function of q and it goes to zero for large q limit. The temperature and the charge density dependence of Seebeck coefficient is drawn in figure 16.

In figure 16(a) and (b), the overall structure of the Seebeck coefficient is similar for both cases except near charge neutrality point. In the absence of magnetic impurity, The behaviors of the Seebeck coefficient is linear in q at charge neutrality point. But if we put magnetic impurity, step-like behavior appears near charge neutrality point for large value of q_{χ} . The Seebeck coefficient near the charge neutrality point at zero temperature is

$$S\Big|_{q\ll 1, T=0} \sim \frac{4\pi}{\beta^2 (1+36q_\chi^2 \gamma^2)} q + \mathcal{O}(q^3).$$
(3.43)

The charge density and the temperature dependences of the Nernst signal are drawn in figure 17.

Nernst signal is proportional to θ from (3.42), therefore it vanishes at zero magnetic impurity case. But in the presence of magnetic impurity, there are maximum and minimum at finite density and temperature as shown in figure 17(b). As we change the ratio of the magnetic impurity γ or interaction strength q_{χ} , the position of the maximum and minimum also changes. Figure 18 shows γ and q_{χ} dependence of the position of the maximum and minimum of Nernst signal.

Numerical study shows that as we increase β , overall shape of Seebeck coefficient is broaden quickly while system does not depends on q_{χ} very much. On the other hand, the Nernst signal has maximum value at finite density and temperature as shown in figure 17(b).



Figure 17. The temperature and the charge density dependence of Seebeck coefficient without (a) with (b) magnetic impurity. We set $\beta^2 = \frac{2747}{(\mu m)^2}$ and $v_F = 7.5 \times 10^4 m/s$.



Figure 18. (a) Maximum value Nernst signal for different γ and q_{χ} . (b) Position of maximum and minimum of Nernst signal.

Increasing β makes the overall shape broaden similar to the Seebeck coefficient. If we increase q_{χ} , overall shape is similar but the position of maximum moves to large q and large T direction and height increases.

3.3 Graphical predictions for Bi₂Se₃

So far we plotted our results with set of parameters such that interesting features appear. For the future experiment, however, it will be more useful to use the parameters which was used to fit magneto conductivity. Here we redraw all the figures such that all the figures can be compared with the data of Bi_2Se_3 . See figures 19–22. One caution is that we set chemical potential zero. Individual material sample can have finite chemical potential for various reason. Gating and impurity doping can bring finite charge and chemical density, for example.

4 Conclusion

In this paper we set up a model for a Dirac material where magnetic impurity is coupled with massless degree of freedom in a time reversal symmetry breaking way. We targeted mainly for the surface states of TI with magnetic impurity doped. However we expect that the model may have more general validity.



Figure 19. The temperature and the magnetic field dependence of (a) parallel, (b) transverse magnetoconductivity. In this subsection we used $v_F = c/3000, \beta = 1.82, q_{\chi}\gamma = 5.6$ in all the figures. The green colored island corresponds to the region where it is expected that our theory might be valid.



Figure 20. The temperature and the magnetic field dependence of (a) parallel, (b) transverse magnetoresitivity.



Figure 21. The temperature and the magnetic field dependence of (a) parallel, (b) transverse thermal conductivity.



Figure 22. The temperature and the magnetic field dependence of (a) Seebeck coefficient, (b) Nernst signal.

From the experience so far, we can mention a few general aspects of Dirac materials. For undoped or weakly doped TI, one normally sees a sharp peak, which is the characteristic of weak anti-localization. We, however, expect that if we can set the fermi surface near Dirac point by gating, we will see the disappearance of the sharp peak as we move down the fermi surface. We also expect that the transition behavior from WAL \rightarrow WL in the medium doping is universal so that magneto-conductivity of all two dimensional Dirac material with broken TRS can be described by our formula, which is independent of the detail of the system. For $\operatorname{Cr}_x \operatorname{Bi}_{2-x} \operatorname{Te}_3$ with x = 0.1 where the system in our picture is strongly interacting for $T \geq 2K$, we expect that ARPES data will show fuzzy density of state (DOS). DOS will be non-zero in the region between dispersion curves, where quasiparticle case would show empty DOS leading to the gap or pseudo gap. Currently we are studying these effects using fermion two point functions. For general Dirac material, we can say that near Dirac point there should be large violation of Wiedemann-Franz Law just like graphene.

Finally we mention some of the future projects. In this paper, we studied the zero charge case mostly. Non-zero charge parameter will be discussed in follow-up paper in relation with the hysteresis of magnetization and various other physical observables. The graphene has even number of Dirac cones, weak spin-orbit interaction and therefore mechanism for WL/WAL is different from the one analyzed here. Different interaction term is necessary. Five dimensional extension of this work will be related to the study of Weyl semi-metal. It is also interesting to study coupling of impurity density with $R \wedge R$ as well as $F \wedge F$. Because of such differences, we need to find other interaction term in holographic model for graphene. It is also interesting to classify all possible patterns of interaction that provide the fermion surface gap in the presence of strong electron-electron correlation in our context.

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A Magnetic induction and magnetization

Here, we describe the problem in magnetization of 2+1 and possible resolution.

- 1. In 2+1 system, we can not add M and H since two have different mass dimensions: [M] = 1, [H] = 2. Therefore although we can theoretically calculate it as a conjugacy of H, we can not compare it with experiment.
- 2. To fix the problem, we can either uplift to or embed our 2+1 dimensional system into 3+1 one. Uplifting is considering a 3+1 system with translational invariance whose dimensionally reduction is the original 2+1 system. Embedding is to consider the 2+1 dimensional system as a thin film in 3+1 with a small thickness L.
- 3. In any case we should have $M_{3+1} = M_{2+1}/L$ for some length scale L. For the thin film case, L is the thickness of the film. For uplifting case, L is inverse temperature if that is the only scale. If we turn on H and q, r_0 is the most natural scale parameter to enter since the energy density $\sim r_0^d$ in any dimension.
- 4. For topological insulator's surface, we do not have a physical scale L so its better to uplift it. We can argue that there is a unique choice of L used in raising $M_{2+1} = LM_{3+1}$: M_{2+1} begins with a term $-\frac{H}{r_0} + \cdots$, at the zero temperature and zero charge and zero impurity limit. That is, if we consider the dyonic black hole, then $r_0 = (12)^{-1/4}\sqrt{H}$ so that $M_{2+1} = -((12)^{1/4})\sqrt{H}$. To find magnetic induction, we need to uplift or embed it to 3+1 as we discussed it by multipying 1/L factor for some length scale L. Then for r_0 independent L, the magnetic induction becomes

$$B = H + M = H - ((12)^{1/4}/L)\sqrt{H} < 0, \quad \text{for} \quad H < \sqrt{12}/L.$$
 (A.1)

This is more than perfect magnetization, which is nonsense. Similar but more complicated argument hold for finite density and temperature. Therefore L should be r_0 dependent.

5. The only way to fix this problem is to choose $1/L = c.r_0$ with some constant c. Now If we seek the energy density such that $M_{3+1} = M_{2+1}/L = -\frac{\partial \varepsilon_{3+1}}{\partial H}$, then we get $\varepsilon_{3+1} = \frac{c}{2}H^2 + \cdots$. Only when we take c = 1, the fist term is identified with energy density of the magnetic field at the vacuum which should be subtracted from the total energy density when we consider the response of the system to the applied magnetic field H [43]. Notice that the uplifting procedure give us a unique way



Figure 23. (a) Temperature dependence of M_0 with for different density. (b) Density dependence of M_0 with different temperature. In both case, we set $q_{\chi} = 1$, $\alpha = 0$ and $\lambda = 2$.

to fix the problem by giving us a natural reasoning to subtract unphysical part of magnetization by attributing it as the derivative of vacuum energy density. Similar description was done in completely condensed matter context [29, 44].

In short, in the absence of a canonical thickness of the surface of TI, we need to uplift to solve the dimensionality problem and to make sure vacuum does not have any non-zero magnetization, we need to redefine the free energy by subtracting the energy of background magnetic field in the absence of the matter,

$$F = r_0 \epsilon - \frac{H^2}{2}, \qquad M \equiv -\frac{\partial F}{\partial H} = \frac{q\theta}{3} - \frac{\theta^2 H}{5}.$$
 (A.2)

The magnetic induction now is identified as

$$B = H + M = H + \frac{q\theta}{3} - \frac{\theta^2 H}{5}.$$
 (A.3)

In the case of the dyonic black hole, $\theta = 0$, and we have M = 0 and hence B = H.

Once we re-define magnetization M, we can study several properties of it. First, the system is still ferromagnetic if charge density is nonzero, that is the magnetization (A.2) has finite value M_0 even in the absence of the external magnetic field:

$$M_0 = \frac{q\theta}{3} = \frac{q\lambda^2 q_{\chi}}{3r_0^2},\tag{A.4}$$

Figure 23(a) shows temperature dependence of M_0 . At zero temperature it becomes

$$M_0(T=0) = \frac{4q\lambda^2 q_{\chi}}{\alpha^2 + \lambda^2 + \sqrt{12q^2 + \alpha^2 + \lambda^2}},$$
 (A.5)

which is proportional to $q\lambda^2 q_{\chi}$ for small λ limit and to qq_{χ} for large λ limit. On the other hand, at high temperature, it suppressed by $1/T^2$ because $r_0 \sim T$ in this regime.

Figure 23(b) shows density dependence of M_0 for different temperature. For large value of q, $r_0 \sim \sqrt{q}$ and the magnetization is saturated to a finite value

$$M_0(q=\infty) \sim \lambda^2 q_{\chi}.$$
 (A.6)



Figure 24. (a) The external magnetic field dependence of the magnetization for different temperature. We fix $\lambda = 2$, $\alpha = 0$, $q_{\chi}=10$ and q = 0. (b) Temperature dependence of the magnetic susceptibility at H = 0 with different value of q_{χ} with $\lambda = 2$, $\alpha = 0$ and q = 0.

We interpret λ^2 as the magnetic impurity density and q_{χ} as the strength of coupling for each magnetic impurity.

Second, in the presence of the external magnetic field, the coefficient of H in the magnetization M in (A.2) is alway negative, so that the system is diamagnetic. At finite magnetic field, the horizon radius (2.11) is a function of the external parameters T, H, β and q_{χ} . The results are shown in figure 24(a). These diamagnetic behavior is similar to the experimental data of certain type of graphite sample [45, 46].

Third, once we obtain the magnetization, we can also calculate the magnetic susceptibility by taking derivative of the magnetization with respect to the external magnetic field;

$$\chi_m \equiv \frac{\delta M}{\delta H} = -\frac{\theta^2}{5} + \frac{4\theta(6H\theta - 5q)(H\theta^2 + H - q\theta)}{15\left\{(3 + 7\theta^2)H^2 - 10Hq\theta + 3q^2 + 2r_0^2(6r_0^2 + \alpha^2 + \lambda^2)\right\}}, \quad (A.7)$$

where second term comes from the variation of r_0 in (A.2). Figure 24(b) shows temperature dependence of the magnetic susceptibility with different value of q_{χ} at zero magnetic field.

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